

Period analysis of the H α equivalent width in P Cygni

The large data set of H α equivalent width measurements (almost 23 years) enables an analysis of its (quasi)-periodic behavior (see fig. below). The data set is a compilation of professional data (Markova & Scuderi) and amateur data, gotten with partly very different spectrographs.

The Scargle-periodogram shows a main period of approx. 1200 days. This is fairly close to the period of the visual brightness of 1300 days, which has been often assumed / discussed in the past, and confirms a certain connection between both, as it has been found in our international campaign "Photometry and Spectroscopy" and has been published in [AAVSO newsletter](#). The other small periods between 500 to 900 days are caused probably by sampling effects (alias periods).

Fig. 1: H α -EW time behavior

Fig. 2: Periodogram

Fig. 3: Phase plot

